







Tutorial on Nonlinear Optics

Robert W. Boyd

Department of Physics and Max-Planck Centre for Extreme and Quantum Photonics University of Ottawa

The visuals of this talk are available at boydnlo.ca/presentations/

Presented at the University of Ottawa Extreme and Quantum Photonics Summer School, June 18-22, 2018.



It is good fundamental physics.

It leads to important applications.

It is a lot of fun.

Demonstrate these features with examples in remainder of talk.

Tutorial on Nonlinear Optics

- 1. What is Nonlinear Optics?
- 2. Coupled Wave Equations and Harmonic Generation
- 3. Mechanisms of Optical Nonlinearity
- 4. Self-Action Effects in Nonlinear Optics
- 5. NLO of Epsilon-Near Zero Materials
- 6. Local-Field Effects in Nonlinear Optics
- 7. Composite Materials and Metamaterials for NLO
- 8. Slow and Fast Light
- 9. Spontaneous and Stimulated Light Scattering

1. What is Nonlinear Optics?

Nonlinear Optics and Light-by-Light Scattering



The elementary process of light-by-light scattering has never been observed in vacuum, but is readily observed using the nonlinear response of material systems.

Nonlinear material is fluorescein-doped boric acid glass (FBAG)

 n_2 (FBAG) $\approx 10^{14} n_2$ (silica) [But very slow response!]

M. A. Kramer, W. R. Tompkin, and R. W. Boyd, Phys. Rev. A, 34, 2026, 1986. W. R. Tompkin, M. S. Malcuit, and R. W. Boyd, Applied Optics 29, 3921, 1990.



Nonlinear Optics

THIRD EDITION





Robert W. Boyd



Simple Formulation of the Theory of Nonlinear Optics

$$P = \chi^{(1)}E + \chi^{(2)}E^2 + \chi^{(3)}E^3 + \dots$$

Here *P* is the induced dipole moment per unit volume and E is the field amplitude

 $\chi^{(1)}$ describes linear optics, e.g., how lenses work: ()

 $\chi^{(2)}$ describes second-order effects, e.g., second-harmonic generation (SHG)

 $\chi^{(3)}$ describes third-order effects such as third-harmonic generation, four-wave mixing, and the intensity dependence of the index of refraction.



Some Fundamental Nonlinear Optical Processes: I



Dolgaleva, Lepeshkin, and Boyd

Some Fundamental Nonlinear Optical Processes: II



Some Fundamental Nonlinear Optical Processes: III



Optical Parametric Oscillation



Parametric Downconversion: A Source of Entangled Photons



 ω_{i}

The signal and idler photons are entangled in:

- (a) polarization
- (b) time and energy
- (c) position and transverse momentum
- (d) angular position and orbital angular momentum

Entanglement is important for:

- (a) Fundamental tests of QM (e.g., nonlocality)
- (a) Quantum technologies (e.g., secure communications)

Some Fundamental Nonlinear Optical Processes: III

Third-Harmonic Generation



Some Fundamental Nonlinear Optical Processes: IV

Intensity-Dependent Index of Refraction

• Single beam: self-phase modulation



• Two beam: cross-phase modulation



The refractive index is given by

 $n = n_0 + n_2 I$ where $n_2 = \frac{3}{4n_0^2\epsilon_0 c}\chi^{(3)}$

Role of Material Symmetry in Nonlinear Optics

 $\chi^{(2)}$ vanishes identically for a material possessing a center of inversion symmetry (a centrosymmetric medium).





centrosymmetric medium





2. Coupled Wave Equations and Harmonic Generation

We have assumed that the pump wave E_1 at frequency ω is undepleted by the nonlinear interaction. We take A_2 to be a function of zto allow the second harmonic wave to grow with z. We also set

 $\tilde{P}_2(t) = P_2 e^{-i2\omega t}$ where $P_2 = \epsilon_0 \chi^{(2)} E_1^2 = \epsilon_0 \chi^{(2)} A_1^2 e^{i2k_1 z}$ (3) The generation of the wave at 2ω is governed by the wave equation

$$\nabla^2 \tilde{E}_2 - \frac{n^2}{c^2} \frac{\partial^2 \tilde{E}_2}{\partial t^2} = \frac{1}{\epsilon_0 c^2} \frac{\partial^2 \tilde{P}_2}{\partial t^2}.$$
 (4)

Note that the first term in the wave equation is given by

$$\nabla^{2}\tilde{E}_{2} = \left[\frac{\partial^{2}\tilde{A}_{2}}{\partial z^{2}} + 2ik_{2}\frac{\partial\tilde{A}_{2}}{\partial z} - k_{2}^{2}A_{2}\right]e^{i(k_{2}z-2\omega t)}$$
(5)
$$\approx \left[2ik_{2}\frac{\partial\tilde{A}_{2}}{\partial z} - k_{2}^{2}A_{2}\right]e^{i(k_{2}z-2\omega t)}$$
(6)

The second form is the *slowly varying amplitude approximation*.

Note also that

$$\frac{\partial^2 \tilde{A}_2}{\partial t^2} = -(2\omega)^2 A_2 e^{i(k_2 z - 2\omega t)} \quad \frac{\partial^2 \tilde{P}_2}{\partial t^2} = -(2\omega)^2 P_2 e^{i(2k_1 z - 2\omega t)} \quad (7)$$

By combining the above equations, we obtain

$$2ik_2\frac{dA_2}{dz} = \frac{-4\omega^2}{c^2}\chi^{(2)}A_1^2e^{i\Delta k\,z} \quad \text{where} \quad \Delta k = 2k_1 - k_2. \tag{8}$$

The quantity Δk is known as the phase (or wavevector) mismatch factor, and it is crucially important in determining the efficiency of nonlinear optical interactions

Treatment of Second-Harmonic Generation – III

For the case $\Delta k = 0$, Eq. (8) becomes

$$2ik_2 \frac{dA_2}{dz} = \frac{-4\omega^2}{c^2} \chi^{(2)} A_1^2 \tag{9}$$

with solution evaluated at z = L of

$$A_2(L) = \frac{2i\omega}{n_2c}\chi^{(2)}A_1^2L \quad \text{or} \quad |A_2(L)|^2 = \frac{4\omega^2}{n_2^2c^2}[\chi^{(2)}]^2|A_1|^4L^2.$$
(10)

Note that the SHG intensity scales as the square of the input intensity and also as the square of the length L of the crystal.



dipole emitter

phased array of dipoles

Treatment of Second-Harmonic Generation – IV

For the general case of $\Delta k \neq 0$, Eq. (8) can still be solved to yield $|A_2(L)|^2 = \frac{4\omega^2}{n_2^2 c^2} [\chi^{(2)}]^2 |A_1|^4 L^2 \operatorname{sinc}^2(\Delta k L/2)$ (11)

Note that $\Delta k L$ must be kept much smaller than π radians in order for efficient SHG to occur.



Nonlinear Optical Microscopy

An important application of harmonic generation is nonlinear microscopy. . .

Microscopy based on second-harmonic generation in the configuration of a confocal microscope and excited by femtosecond laser pulses was introduced by Curley et al. (1992). Also, harmonic-generation microscopy can be used to form images of transparent (phase) objects, because the phase matching condition of nonlinear optics depends sensitively on the refractive index variation within the sample being imaged (Muller et al., 1998).

Boyd, NLO, Subsection 2.7.1

Caution!

Curley et al., not Curly et al.



How to Achieve Phase Matching: Birefringence Phase Matching

The phase matching condition $\Delta k = 0$ requires that

$$\frac{n_1\omega_1}{c} + \frac{n_2\omega_2}{c} = \frac{n_3\omega_3}{c} \quad \text{where} \quad \omega_1 + \omega_2 = \omega_3$$

These conditions are incompatible in an isotropic dispersive material. However, for a birefringent material phase matching can be achieved.



Midwinter and Warner showed that there are two ways to achieve phase matching:

	Positive uniaxial	Negative uniaxial
	$(n_e > n_0)$	$(n_e < n_0)$
Type I Type II	$n_{3}^{0}\omega_{3} = n_{1}^{e}\omega_{1} + n_{2}^{e}\omega_{2}$ $n_{3}^{0}\omega_{3} = n_{1}^{o}\omega_{1} + n_{2}^{e}\omega_{2}$	$n_3^e \omega_3 = n_1^o \omega_1 + n_2^o \omega_2$ $n_3^e \omega_3 = n_1^e \omega_1 + n_2^o \omega_2$



Sign of $\chi^{(2)}$ is periodically inverted to prevent reverse power flow.



z / L_{coh}

Nonlinear Optics with Focused Laser Beams
(Real experiments don't use infinite plane waves)
Focus beam to increase laser intensity.
What is optimum degree of focusing 7
(too loose) (just right) (too tight)
Trade off between peak intensity P/TWo² and
length of interaction region ~ b.

$$2W_0$$

 $b = 2\pi W_0^2$
 $k = 2\pi W_0^2$
For SHG, Boyd and Kleinman (1968) showed
that maximum power conversion occurs for
 $b = L/2.84$
 $\Delta k = 3.2/L$ Gouy phase
shift!
and is given by
 $\frac{P(w)}{P(w)} = 1.068 \left[\frac{128 \pi^2 w_1^3 deff}{c^4 n_1 n_2}\right] P(w)$

-

Additional Studies of Wave Propagation Effects



3. Mechanisms of Optical Nonlinearity

Typical Values of the Nonlinear Refractive Index

Mechanism	${n_2}^a$ $(\mathrm{cm}^2/\mathrm{W})$	$\chi^{(3)}_{1111} \ (m^2/V^2)$	Response time (sec)
Electronic polarization	10^{-16}	10^{-22}	10^{-15}
Molecular orientation	10^{-14}	10^{-20}	10^{-12}
Electrostriction	10^{-14}	10^{-20}	10^{-9}
Saturated atomic absorption	10^{-10}	10^{-16}	10^{-8}
Thermal effects	10^{-6}	10^{-12}	10^{-3}
Photorefractive $effect^b$	(large)	(large)	(intensity-dependent)

 a For linearly polarized light.

^b The photorefractive effect often leads to a very strong nonlinear response. This response usually cannot be described in terms of a $\chi^{(3)}$ (or an n_2) nonlinear susceptibility, because the nonlinear polarization does not depend on the applied field strength in the same manner as the other mechanisms listed.

Quantum Mechanical Origin of the Nonlinear Optical Suesceptibility

$$\begin{split} \chi_{ijk}^{(2)}(\omega_{p} + \omega_{q}, \omega_{q}, \omega_{p}) \\ &= \frac{N}{2 \epsilon_{0} \hbar^{2}} \sum_{lmn} \rho_{ll}^{(0)} \bigg\{ \frac{\mu_{ln}^{i} \mu_{nm}^{j} \mu_{ml}^{k}}{[(\omega_{nl} - \omega_{p} - \omega_{q}) - i\gamma_{nl}][(\omega_{ml} - \omega_{p}) - i\gamma_{ml}]} \\ &+ \frac{\mu_{ln}^{i} \mu_{nm}^{k} \mu_{ml}^{j}}{[(\omega_{nl} - \omega_{p} - \omega_{q}) - i\gamma_{nl}][(\omega_{ml} - \omega_{q}) - i\gamma_{ml}]} \quad (a_{2}) \\ &+ \frac{\mu_{ln}^{k} \mu_{nm}^{i} \mu_{ml}^{j}}{[(\omega_{mn} - \omega_{p} - \omega_{q}) - i\gamma_{mn}][(\omega_{nl} + \omega_{p}) + i\gamma_{nl}]} \quad (a'_{1}) \\ &+ \frac{\mu_{ln}^{j} \mu_{nm}^{i} \mu_{ml}^{k}}{[(\omega_{mn} - \omega_{p} - \omega_{q}) - i\gamma_{mn}][(\omega_{nl} + \omega_{q}) + i\gamma_{nl}]} \quad (a'_{2}) \\ &+ \frac{\mu_{ln}^{j} \mu_{nm}^{i} \mu_{ml}^{k}}{[(\omega_{nm} + \omega_{p} + \omega_{q}) + i\gamma_{nm}][(\omega_{ml} - \omega_{p}) - i\gamma_{ml}]} \quad (b_{1}) \\ &+ \frac{\mu_{ln}^{k} \mu_{nm}^{i} \mu_{ml}^{j}}{[(\omega_{nm} + \omega_{p} + \omega_{q}) + i\gamma_{nm}][(\omega_{ml} - \omega_{q}) - i\gamma_{ml}]} \quad (b_{2}) \\ &- \frac{\mu_{ln}^{k} \mu_{nm}^{i} \mu_{ml}^{j}}{[(\omega_{nm} + \omega_{p} + \omega_{q}) + i\gamma_{nm}][(\omega_{ml} - \omega_{q}) - i\gamma_{ml}]} \\ &+ \frac{\mu_{ln}^{k} \mu_{nm}^{i} \mu_{ml}^{j}}{[(\omega_{nm} + \omega_{p} + \omega_{q}) + i\gamma_{nm}][(\omega_{ml} - \omega_{q}) - i\gamma_{ml}]} \\ &+ \frac{\mu_{ln}^{k} \mu_{nm}^{i} \mu_{ml}^{j}}{[(\omega_{nm} + \omega_{p} + \omega_{q}) + i\gamma_{nm}][(\omega_{ml} - \omega_{q}) - i\gamma_{ml}]} \\ &+ \frac{\mu_{ln}^{k} \mu_{nm}^{i} \mu_{ml}^{j}}{[(\omega_{nm} + \omega_{p} + \omega_{q}) + i\gamma_{nm}][(\omega_{ml} - \omega_{q}) - i\gamma_{ml}]} \\ &+ \frac{\mu_{ln}^{k} \mu_{nm}^{i} \mu_{ml}^{j}}{[(\omega_{nm} + \omega_{p} + \omega_{q}) + i\gamma_{nm}][(\omega_{ml} - \omega_{q}) - i\gamma_{ml}]} \\ &+ \frac{\mu_{ln}^{k} \mu_{nm}^{i} \mu_{ml}^{j}}{[(\omega_{nm} + \omega_{p} + \omega_{q}) + i\gamma_{nm}][(\omega_{ml} - \omega_{q}) - i\gamma_{ml}]} \\ &+ \frac{\mu_{ln}^{k} \mu_{nm}^{i} \mu_{ml}^{j}}{[(\omega_{nm} + \omega_{p} + \omega_{q}) + i\gamma_{nm}][(\omega_{ml} - \omega_{q}) - i\gamma_{ml}]} \\ &+ \frac{\mu_{ln}^{k} \mu_{nm}^{i} \mu_{ml}^{j}}{[(\omega_{nm} + \omega_{p} + \omega_{q}) + i\gamma_{nm}][(\omega_{ml} - \omega_{q}) - i\gamma_{ml}]} \\ &+ \frac{\mu_{ln}^{k} \mu_{mn}^{i} \mu_{ml}^{j}}{[(\omega_{nm} + \omega_{p} + \omega_{q}) + i\gamma_{mm}][(\omega_{ml} - \omega_{q}) - i\gamma_{ml}]} \\ &+ \frac{\mu_{ln}^{k} \mu_{mn}^{i} \mu_{ml}^{j}}{[(\omega_{nm} + \omega_{p} + \omega_{q}) + i\gamma_{mm}]} \\ &+ \frac{\mu_{ln}^{k} \mu_{mn}^{i} \mu_{mn}^{j}}{[(\omega_{nm} + \omega_{p} + \omega_{q}) + i\gamma_{mm}]} \\ &+ \frac{\mu_{ln}^{k} \mu_{mn}^{i} \mu_{mn}^{j}}{[(\omega_{mn} + \omega_{p} + \omega_{mn}) + i\gamma_{mm}]} \\ &+ \frac{\mu_{ln}^{k} \mu_{mn}^{i} \mu_{mn}^{j}}{[(\omega_{mn} + \omega_{p} +$$

$$+ \frac{\mu_{ln}^{*}\mu_{nm}^{*}\mu_{ml}^{*}}{[(\omega_{ml} + \omega_{p} + \omega_{q}) + i\gamma_{ml}][(\omega_{nl} + \omega_{p}) + i\gamma_{nl}]} \qquad (b_{1}')$$

$$+ \frac{\mu_{ln}^{J}\mu_{nm}^{k}\mu_{ml}^{i}}{[(\omega_{ml}+\omega_{p}+\omega_{q})+i\gamma_{ml}][(\omega_{nl}+\omega_{q})+i\gamma_{nl}]} \bigg\}$$
 (b₂)





Nonresonant Electronic Nonlinearities Estimate size:

$$P = \chi^{(1)}E + \chi^{(3)}E^{3}$$

nonlinear term becomes comparable to linear term when

$$= \frac{E}{E_{at}} = \frac{e}{a_{s}^{2}} = \frac{B_{ohr}}{radius}$$

$$= \frac{\chi^{(3)}}{E_{at}^{2}} \approx \frac{\chi^{(0)}}{E_{at}^{2}} \approx \frac{1}{E_{at}^{2}} = 3.8 \times 10^{-24} \text{ m}^{2}/\sqrt{2}$$

Must generalize the Lorentz model of atom to allow a nonlinearity in restoring force.





How to Increase
$$\chi^{(3)}$$

Use a conjugated polymer

 $\chi \chi \chi \chi$
lettertron wavefunction
delocalized over length
of polymer
linear polarizability $\chi \sim L^3$
nonlinear hyperpolarizability $\chi \sim L^5$
typically $\chi^{(3)} \simeq 10^{-11} escufor conjugated polymer$

•



.

4. Self-Action Effects in Nonlinear Optics

Self Action Effects in Nonlinear Optics

Self-action effects: light beam modifies its own propagation

self focusing







small-scale filamentation



Why Care About Self-Focusing and Filamentation?

- Optical switching
- Laser modelocking
- Directed energy

 prevent filamentation
 controlled self focusing

EFFECTS OF THE GRADIENT OF A STRONG ELECTROMAGNETIC BEAM ON ELECTRONS AND ATOMS

G. A. ASKAR'YAN

P. N. Lebedev Physics Institute, Academy of Sciences, U.S.S.R.

Submitted to JETP editor December 22, 1961

J. Exptl. Theoret. Phys. (U.S.S.R.) 42, 1567-1570 (June, 1962)

It is shown that the transverse inhomogeneity of a strong electromagnetic beam can exert a strong effect on the electrons and atoms of a medium. Thus, if the frequency exceeds the natural frequency of the electron oscillations (in a plasma or in atoms), then the electrons or atoms will be forced out of the beam field. At subresonance frequencies, the particles will be pulled in, the force being especially large at resonance. It is noted that this effect can create either a rarefaction or a compression in the beam and at the focus of the radiation, maintain a pressure gradient near an opening from an evacuated vessel to the atmosphere, and create a channel for the passage of charged particles in the medium.

It is shown that the strong thermal ionizing and separating effects of the ray on the medium can be used to set up waveguide propagation conditions and to eliminate divergence of the beam (self-focusing). It is noted that hollow beams can give rise to directional flow and ejection of the plasma along the beam axis for plasma transport and creation of plasma current conductors. The possibilities of accelerating and heating plasma electrons by a modulated beam are indicated.
Prediction of Self Trapping



Z-Scan Measurement of X(3) Laser aperture Measures NL change in refraction $(\operatorname{Re} X^{(3)})$ Z Laser No aperture Measures NL change (or open aperture) in absorption. $(\operatorname{Im}\chi^{(3)})$

ź



For closed aperture z-scan

 $\Delta T_{pv} = 0.406 \ \Phi^{\text{NL}} \label{eq:Tpv}$ where

 $\Phi^{\text{NL}} = n_2 (\omega/c) I_0 L$

M. Sheik-Bahae et al., IEEE J. Quantum Electron. 26 760 (1990). Predicted by Bespalov and Talanov (1966)

Exponential growth of wavefront imperfections by four-wave mixing processes



Optical Solitons

Field distributions that propagate without change of form

Temporal solitons (nonlinearity balances gvd)

$$\frac{\partial \tilde{A}_s}{\partial z} + \frac{1}{2}ik_2\frac{\partial^2 \tilde{A}_s}{\partial \tau^2} = i\gamma |\tilde{A}_s|^2 \tilde{A}_s.$$

1973: Hasegawa & Tappert 1980: Mollenauer, Stolen, Gordon



$$2ik_0\frac{\partial A}{\partial z} + \frac{\partial^2 A}{\partial x^2} = -3\chi^{(3)} \frac{\omega^2}{c^2} |A|^2 A$$

1964: Garmire, Chiao, Townes
1974: Ashkin and Bjorkholm (Na)
1985: Barthelemy, Froehly (CS2)
1991: Aitchison et al. (planar glass waveguide
1992: Segev, (photorefractive)



-10 0 100

Solitons and self-focussing in Ti:Sapphire

42 OPTICS LETTERS / Vol. 16, No. 1 / January 1, 1991

60-fsec pulse generation from a self-mode-locked Ti:sapphire laser

D. E. Spence, P. N. Kean, and W. Sibbett

J. F. Allen Physics Research Laboratories, Department of Physics and Astronomy, University of St. Andrews, North Haugh, St. Andrews, Fife, KY16 9SS, Scotland



Diffraction-management controls the spatial selffocussing

Dispersion-management controls the temporal self-focussing

Self-Focusing Can Produce Unusual Beam Patterns

Pattern depends sensitively upon initial conditions

• Conical emission Harter et al., PRL 46, 1192 (1981)



• Multiple ring patterns Kauranen et al, Opt. Lett. 16, 943, 1991;



• Honeycomb pattern formation Bennink et al., PRL 88, 113901 2002.



• Loss of spatial coherence Schweinsberg et al., Phys. Rev. A 84, 053837 (2011).



Summary

- Even more than 50 years after their inceptions, self-focusing and filamentation remain fascinating topics for investigation.
- If you want to learn more:



5. NLO of Epsilon-Near Zero Materials

- Physics of Epsilon-Near-Zero (ENZ) Materials
- Huge NLO Response of ENZ Materials and Metastructures
- Non-perturbative nature of the NLO Response (usual power series do not converge)

With Special Thanks To:

M. Zahirul Alam, Orad Reshef, Enno Giese, and Jeremy Upham, University of Ottawa Israel De Leon, Tecnologico de Monterrey, Mexico Sebastian Schulz, Cork Institute of Technology, Ireland

Physics of Epsilon-Near-Zero (ENZ) Materials

- ENZ materials possess exotic electromagnetic properties Silveirinha, Engheta, Phys. Rev. Lett. 97, 157403, 2006.
- If the dielectric permittivity ε is nearly zero, then refractive index $n = \operatorname{sqrt}(\varepsilon)$ is nearly zero.

Thus $v_{\text{phase}} = c / n$ is nearly infinite

 $\lambda = \lambda_{vac} / n$ is nearly infinite

Light oscillates in time but not in space; everyhing is in phase Light "oscillates" but does not "propagate."

• Radiative processes are modified in ENZ materials

Einstein A coefficient (spontaneous emission lifetime = 1/A)

 $A = n A_{vac}$

We can control (inhibit!) spontaneous emission!

Einstein *B* coefficient

Stimulated emission rate = *B* times EM field energy density

 $B = B_{\rm vac} / n^2$

Optical gain is very large!

Einstein, Physikalische Zeitschrift 18, 121 (1917). Milonni, Journal of Modern Optics 42, 1991 (1995).

Physics of Epsilon-Near-Zero (ENZ) Materials -- More

Snell's law leads to intriguing predictions

$$n_1 \sin \theta_1 = n_2 \sin \theta_2$$

• Light always leaves perpendicular to surface of ENZ material!



Y. Li, et al., Nat. Photonics 9, 738, 2015; D. I. Vulis, et al., Opt. Express 25, 12381, 2017.

• Thus light can enter an ENZ material only at normal incidence!



Y. Li, et al., Nat. Photonics 9, 738, 2015.

Some Consequences of ENZ Behaviour - 1

• Funny lenses



A. Alù et al., Phys. Rev. B 75, 155410, 2007; X.-T. He, ACS Photonics, 3, 2262, 2016.

• Large-area single-transverse-mode surface-emitting lasers

J. Bravo-Abad et al., Proc. Natl. Acad. Sci. USA 109, 976, 2012.

• No Fabry-Perot interference



O. Reshef et al., ACS Photonics 4, 2385, 2017.

Some Consequences of ENZ Behaviour - 2

• Super-coupling (of waveguides)



M. G. Silveirinha and N. Engheta, Phys. Rev. B 76, 245109, 2007; B. Edwards et al., Phys. Rev. Lett. 100, 033903, 2008.

• Large evanescent tails for waveguide coupling

transverse profile of upper waveguide extends to lower waveguide for any distance

[•] dielectric waveguide

Automatic phase matching of NLO processes

Recall that $k = n \omega / c$ vanishes in an ENZ medium.

For example, the following 4WM proces is allowed



H. Suchowski et al., Science 342, 1223, 2013.

Some Consequences of ENZ Behaviour - 3

- How is the theory of self-focusing modified?
- Does the theory of Z-scan need to be modified?
- How is the theory of blackbody radiation modified?
- Do we expect very strong superradiance effects?
- More generally, how is any NLO process modified when $n_0 = 0$?

- Metamaterials
 Materials tailor-made to display ENZ behaviour
- Homogeneous materials

All materials display ENZ behaviour at their (reduced) plasma frequency

Recall the Drude formula

$$\epsilon(\omega) = \epsilon_{\infty} - \frac{\omega_p^2}{\omega(\omega + i\gamma)}$$

Note that $\operatorname{Re} \epsilon = 0$ for $\omega = \omega_p / \sqrt{\epsilon_\infty} \equiv \omega_0$.

- Challenge: Obtain low-loss ENZ materials Want Im ϵ as small as possible at the frequency where Re $\epsilon = 0$.
- We are examining a several materials ITO: indium tin oxide AZO: aluminum zinc oxide FTO: fluorine tin oxide

Epsilon-Near-Zero Materials for Nonlinear Optics

- We need materials with a much larger NLO response
- We recently reported a material (indium tin oxide, ITO) with an n_2 value 100 time larger than those previously reported.
- This material utilizes the strong enhancement of the NLO response that occurs in the epsilon-near zero (ENZ) spectral region.

Large optical nonlinearity of indium tin oxide in its epsilon-near-zero region, M. Zahirul Alam, I. De Leon, R. W. Boyd, Science 352, 795 (2016).

Here is the intuition for why the ENZ conditions are of interest in NLO Recall the standard relation between n_2 and $\chi^{(3)}$

$$n_2 = \frac{3\chi^{(3)}}{4\epsilon_0 c \, n_0 \operatorname{Re}(n_0)}$$

Note that for ENZ conditions the denominator becomes very small, leading to a very large value of n_2

ITO is a degenerate semiconductor (so highly doped as to be metal-like).

It has a very large density of free electrons, and a bulk plasma frequency corresponding to a wavelength of approximately 1.24 µm.

Recall the Drude formula

$$\epsilon(\omega) = \epsilon_{\infty} - \frac{\omega_p^2}{\omega(\omega + i\gamma)}$$

Note that $\operatorname{Re} \epsilon = 0$ for $\omega = \omega_p / \sqrt{\epsilon_\infty} \equiv \omega_0$.

The region near ω_0 is known as the epsilon-near-zero (ENZ) region.

There has been great recent interest in studies of ENZ phenomena:

H. Suchowski, K. O'Brien, Z. J. Wong, A. Salandrino, X. Yin, and X. Zhang, Science 342, 1223 (2013).
C. Argyropoulos, P.-Y. Chen, G. D'Aguanno, N. Engheta, and A. Alu, Phys. Rev. B 85, 045129 (2012).
S. Campione, D. de Ceglia, M. A. Vincenti, M. Scalora, and F. Capolino, Phys. Rev. B 87, 035120 (2013).
A. Ciattoni, C. Rizza, and E. Palange, Phys. Rev. A 81,043839 (2010).

Huge Nonlinear Optical Response Measured by Z-scan



- Note that n_2 is positive (self focusing) and β is negative (saturable absorption)
- Both n_2 and nonlinear absorption increase with angle of incidence
- n_2 shows a maximum value of 0.11 cm²/GW = 1.1 × 10⁻¹⁰ cm²/W at 1.25 µm and 60 deg. This value is 2000 times larger than that away from ENZ region.
- n_2 is 3.4 x 10⁵ times larger than that of fused silica n_2 is 200 times larger than that of chalcogenide glass

Beyond the $\chi^{(3)}$ limit



The nonlinear change in refractive index is so large as to change the transmission, absorption, and reflection!

Note that transmission is increased at high intensity.

Here is the refractive index extracted from the above data.

Note that the total nonlinear change in refractive index is $\Delta n = 0.8$.

The absorption decreases at high intensity, allowing a predicted NL phase shift of 0.5 radians.

Nonperturbative Nature of the NLO Response

- 1. The conventional equation $n = n_0 + n_2 I$ is not applicable to ENZ and other low-index materials. The nonlinear response is nonperturbative.
- 2. The nonlinear response can be accurately modeled in the $\chi^{(3)}$ limit by

$$n = \sqrt{n_0^2 + 2n_0 n_2 I}$$

where

$$n_2 = \frac{3\chi^{(3)}}{4n_0 \operatorname{Re}(n_0)\epsilon_0 c}.$$

and

$$I = 2\operatorname{Re}(n_0)\epsilon_0 c|E|^2$$

3. More generally, the intensity dependent refractive index can be described by

$$n = \sqrt{\epsilon^{(1)} + 3\chi^{(3)}|E|^2 + 10\chi^{(5)}|E|^4 + \cdots}$$

- 1. The conventional equation $n = n_0 + n_2 I$ is not applicable to ENZ and other low-index materials. The nonlinear response is nonperturbative.
- 2. The nonlinear response can be accurately modeled in the $\chi^{(3)}$ limit by

$$n = \sqrt{n_0^2 + 2n_0 n_2 I}$$

where

$$n_2 = \frac{3\chi^{(3)}}{4n_0 \operatorname{Re}(n_0)\epsilon_0 c}.$$

and

$$I = 2\operatorname{Re}(n_0)\epsilon_0 c|E|^2$$

3. More generally, the intensity dependent refractive index can be described by

$$n = \sqrt{\epsilon^{(1)} + 3\chi^{(3)}|E|^2 + 10\chi^{(5)}|E|^4 + \cdots}$$



6. Local-Field Effects in Nonlinear Optics

Recall the Lorentz-Lorenz Law (linear optics)

$$\chi^{(1)} = \frac{N\alpha}{1 - \frac{4}{3}\pi N\alpha}$$
 or $\frac{\epsilon^{(1)} - 1}{\epsilon^{(1)} + 2} = \frac{4}{3}\pi N\alpha.$



This result follows from the assumption that the field that acts on a representative atom is not the macroscopic Maxwell field but rather the Lorentz local field given by

$$E_{\text{loc}} = E + \frac{4}{3}\pi P$$
 where $P = \chi^{(1)}E$

We can rewrite this result as

$$E_{\text{loc}} = LE$$
 where $L = \frac{\epsilon^{(1)} + 2}{3}$ is the local field factor.

The Lorentz Red Shift

$$\frac{\varepsilon - 1}{\varepsilon + 2} = \frac{4\pi}{3} N \alpha$$

.

$$\alpha = \frac{c f re \lambda_0 e / 4\pi}{W_0 - W - i \gamma}$$

$$\mathcal{E} = 1 + \frac{Nfre \lambda_{o}c}{W_{o} + \Delta W_{L} - W - i\delta}$$

$$\Delta W_{L} = -\frac{1}{3}Nfre \lambda_{o}c$$

$$T$$

See, for instance, H. A. Lorentz, Theory of Electrons, Dover, NY (1952).

Observation of the Lorentz Red Shift



Maki, Malcuit, Sipe, and Boyd, Phys. Rev. Lett. 68, 972 (1991).

For the case of nonlinear optics, Bloembergen (1962, 1965) showed that, for instance,

$$\chi^{(3)}(\omega = \omega + \omega - \omega) = N\gamma^{(3)}|L(\omega)|^2[L(\omega)]^2.$$

where $\gamma^{(3)}$ is the second hyperpolarizability and where

$$L(\omega) = \frac{\epsilon(\omega) + 2}{3}$$

For the typical value n = 2, L = 2, and $L^4 = 16$. Local field effects can be very large in nonlinear optics! But can we tailor them for our benefit?

We have been developing new photonic materials with enhanced NLO response by using composite structures that exploit local field effects.

7. Composite Materials and Metamaterials for NLO

Nanocomposite Materials for Nonlinear Optics



- In each case, scale size of inhomogeneity << optical wavelength
- Thus all optical properties, such as *n* and $\chi^{(3)}$, can be described by effective (volume averaged) values

Recent review: Dolgaleva and Boyd, Advances in Optics and Photonics 4, 1–77 (2012).

Local Field Enhancement of the NLO Response

- Under very general conditions, we can express the NL response as $\chi_{eff}^{(3)} = f L^2 |L|^2 \chi^{(3)}$

where *f* is the volume fraction of nonlinear material and *L* is the local-field factor, which is different for each material geometry.

- Under appropriate conditions, the product $fL^2|L|^2$ can exceed unity.
- For a homogeneous material $L = \frac{\varepsilon + 2}{3}$
- For a spherical particle of dielectric constant ε_m embedded in a host of dielectric constant ε_h $I = \frac{3\varepsilon_h}{1 \frac{3\varepsilon_h}{2}}$

$$L = \frac{3\varepsilon_h}{\varepsilon_m + 2\varepsilon_h}$$

• For a layered geometry with the electric field perpendicular to the plane of the layers, the local field factor for component a is given by

$$L = \frac{\varepsilon_{\rm eff}}{\varepsilon_{\rm a}} \qquad \frac{1}{\varepsilon_{\rm eff}} = \frac{f_{\rm a}}{\varepsilon_{\rm a}} + \frac{f_{\rm b}}{\varepsilon_{\rm b}}$$

Material Systems for Composite NLO Materials

All-dielectric composite materials

Minimum loss, but limited NL response

Metal-dielectric composite materials

Larger loss, but larger NL response

Note that $\chi^{(3)}$ of gold $\approx 10^6 \chi^{(3)}$ of silica glass!

Also, metal-dielectric composites possess surface plasmon resonances, which can further enhance the NL response.

Enhanced NLO Response from Layered Composite Materials

A composite material can display a larger NL response than its constituents!

Alternating layers of TiO₂ and the conjugated polymer PBZT.



 $\nabla \cdot \mathbf{D} = 0$ implies that $(\varepsilon \mathbf{E})_{\perp}$ is continuous.

Measure NL phase shift as a function of angle of incidence.

35% enhancement in $\chi^{(3)}$

Fischer, Boyd, Gehr, Jenekhe, Osaheni, Sipe, and Weller-Brophy, Phys. Rev. Lett. 74, 1871 (1995).



Quadratic EO effect

3.2 times enhancement!

Nelson and Boyd, APL 74 2417 (1999)

$$\frac{\text{Metal} / \underline{\text{Dielectric Composites}}}{\text{Very large local field effects}}$$

$$\frac{\text{Very large local field effects}}{E_{h}} = \frac{3E_{h}}{E_{m}+2E_{h}} E_{o}$$

$$E_{h} = Z E_{o}$$

$$(E_{m} \text{ is negative } !)$$

$$At \text{ resonance}$$

$$Z = \frac{3E_{h}}{E_{m}+2E_{h}} \rightarrow \frac{3E_{h}}{iE_{m}''} \approx (3 + 0 30) i$$

Gold-Doped Glass: A Maxwell-Garnett Composite



Red Glass Caraffe Nurenberg, ca. 1700

Huelsmann Museum, Bielefeld

Developmental Glass, Corning Inc.

gold volume fraction approximately 10⁻⁶ gold particles approximately 10 nm diameter



- Composite materials can possess properties very different from those of their constituents.
- Red color is because the material absorbs very strong in the blue, at the surface plasmon frequency
Accessing the Optical Nonlinearity of Metals with Metal-Dielectric Photonic Crystal Structures

- Metals have very large optical nonlinearities but low transmission
- Low transmission because metals are highly reflecting (not because they are absorbing!)
- Solution: construct metal-dielectric photonic crystal structure (linear properties studied earlier by Bloemer and Scalora)



 $I = 500 \text{ MW/cm}^2$

8. Slow and Fast Light

Controlling the Velocity of Light

"Slow," "Fast" and "Backwards" Light

- Light can be made to go: slow: $v_g \ll c$ (as much as 10⁶ times slower!) fast: $v_g > c$ backwards: v_g negative Here v_g is the group velocity: $v_g = c/n_g$ $n_g = n + \omega (dn/d\omega)$
- Velocity controlled by structural or material resonances





Review article: Boyd and Gauthier, Science 326, 1074 (2009).



Group velocity given by $V_{\overline{3}} = \frac{dW}{dR}$ For $k = \frac{n\omega}{c}$ $\frac{dk}{d\omega} = \frac{1}{c} \left(n + \omega \frac{dn}{d\omega} \right)$

Thus

 $V_{g} = \frac{c}{n + \omega \frac{dn}{d\omega}} \equiv \frac{c}{n_{g}}$

Thus $n_g \neq n$ in a dispersive medium!

Slow Light in a Fiber Bragg Grating (FBG) Structure

(Can describe properties of FBGs by means of analytic expressions)

forward and backward waves are strongly coupled

$$E_{F}(0,t) \longrightarrow D \longrightarrow E_{F}(L,t)$$

$$E_{B}(0,t) \iff E_{F}(L,t)$$





• Enhanced NLO response

Bhat and Sipe showed that the nonlinear coeficient is given by

$$\Gamma = \left(\frac{3 - S^{-2}}{2}\right) S^2 \gamma_0$$

where the slow-down factor $S = n_g/n$

Improved Slow-Light Fiber Bragg Grating (FBG) Structure



H. Wen, M. Terrel, S. Fan and M. Digonnet, IEEE Sensors J. 12, 156-163 (2012).

J. Upham, I. De Leon, D. Grobnic, E. Ma, M.-C. N. Dicaire, S.A. Schulz, S. Murugkar, and R.W. Boyd, Optics Letters 39, 849-852 (2014).

Slow Light Fundamentals: How to Create Slow and Fast Light I Use Isolated Gain or Absorption Resonance



Light speed reduction to 17 metres per second in an ultracold atomic gas

Lene Vestergaard Hau^{*2}, S. E. Harris³, Zachary Dutton^{*2} & Cyrus H. Behroozi^{*}§

* Rowland Institute for Science, 100 Edwin H. Land Boulevard, Cambridge, Massachusetts 02142, USA

² Department of Physics, § Division of Engineering and Applied Sciences, Harvard University, Cambridge, Massachusetts 02138, USA

³ Edward L. Ginzton Laboratory, Stanford University, Stanford, California 94305, USA

NATURE | VOL 397 | 18 FEBRUARY 1999 |www.nature.com

0.8 0.7

0.6

0.5

0.4 0.3

0.2

0.1

1.006

1.004

1.002 1.000

0.998 0.996 0.994

-30

-30

Transmission

Refractive index

а





Note also related work by Chu, Wong, Welch, Scully, Budker, Ketterle, and many others

Slow Light and Fast Light: A Tutorial Review

- 1. Introduction to slow light
- 2. Mechanisms of slow light
- 3. Backwards light and causality
- 4. Slow light spectrometer
- 5. Optical forces
- 6. Fresnel drag
- 7. Other applications

Slow Light in Room-Temperature Solid-State Materials

- Crucial for many real-world applications
- Two examples
 - Slow light *via* stimulated Brillouin scattering (SBS)
 - Slow light *via* coherent population oscillations (CPO)

Slow Light by Stimulated Brillouin Scattering (SBS)



We often think of SBS as a pure gain process, but it also leads to a change in refractive index



The induced time delay is $\Delta T_{\rm d} \approx \frac{G}{\Gamma_{\rm B}}$ where $G = g I_{\rm P} L$ and $\Gamma_{\rm B}$ is the Brillouin linewidth

Okawachi, Bigelow, Sharping, Zhu, Schweinsberg, Gauthier, Boyd, and Gaeta Phys. Rev. Lett. 94, 153902 (2005). Related results reported by Song, González Herráez and Thévenaz, Optics Express 13, 83 (2005).

Slow Light via Coherent Population Oscillations



- profile to give a large $dn/d\omega$
- Ground state population oscillates at beat frequency δ (for $\delta < 1/T_1$).
- Population oscillations lead to decreased probe absorption (by explicit calculation), even though broadening is homogeneous.
- Ultra-slow light ($n_g > 10^6$) observed in ruby and ultra-fast light $(n_g = -4 \times 10^5)$ observed in alexandrite.
- Slow and fast light effects occur at room temperature!

PRL 90,113903(2003); Science, 301, 200 (2003)

Relation of CPO to the Basov Mechanism

 CPO slow light: a strong pump beam creates a narrow transparency window, and strong spectral variation of the refractive index leads to a large group index.



- Basov mechanism: an isolated intense pulse passing through a saturable material experiences a time delay.
 - Assume that $T_{pulse} \ll T_1 = time scale for saturation changes$
 - Then absorption decreases with time during pulse due to saturation



Slow Light in Ruby



Alexandrite Displays both Saturable and Reverse-Saturable Absorption

• Both slow and fast propagation observed in alexandrite



Bigelow, Lepeshkin, and Boyd, Science 301, 200 (2003).

Slow and Fast Light in an Erbium Doped Fiber Amplifier

- Fiber geometry allows long propagation length
- Saturable gain or loss possible depending on pump intensity





Slow Light and Fast Light: A Tutorial Review

- 1. Introduction to slow light
- 2. Mechanisms of slow light
- **3. Backwards light and causality**
- 4. Slow light spectrometer
- 5. Optical forces
- 6. Fresnel drag
- 7. Other applications

Observation of Superluminal and "Backwards" Pulse Propagation

- A strongly counterintuitive phenomenon
- But entirely consistent with established physics
- Predicted by Garrett and McCumber (1970) and Chiao (1993).
- Observed by Gehring, Schweinsberg, Barsi, Kostinski, and Boyd Science 312, 985 2006.





SLO

Observation of Backward Pulse Propagation in an Erbium-Doped-Fiber Optical Amplifier



R.W. Boyd, Science 312, 985 2006.

Causality?

- Superluminal (v_g>c) and backwards (v_g negative) propagation may seem counterintuitive but are fully compatible with causality.
- The group velocity is the velocity at which peak of pulse moves; it is not the "information velocity."
- It is believed that information is carried by points of nonalyticity of a waveform



- broad spectral content at points of discontinuity
- disturbance moves at vacuum speed of light

see, for instance, R.Y. Chiao

How to Reconcile Superluminality with Causality



Stenner, Gauthier, and Neifeld, Nature, 425, 695, 2003.

Bigelow, Lepeshkin, Shin, and Boyd, J. Phys: Condensed Matter, 3117, 2006.

Gauthier and Boyd, Photonics Spectra, p. 82 January 2007.

Slow Light and Fast Light: A Tutorial Review

- 1. Introduction to slow light
- 2. Mechanisms of slow light
- 3. Backwards light and causality
- 4. Slow light spectrometer
- 5. Optical forces
- 6. Fresnel drag
- 7. Other applications

Chip-Scale Spectrometers for Chem-Bio Identification

- Spectroscopy is the standard laboratory procedure for identifying chemical species.
- Can we fabricate miniaturized, chip-scale spectrometers without sustaining a loss in resolution?









Development of Miniaturized, Chip-Scale Spectrometers

Can We Beat the 1/L Resolution Limit of Standard Spectrometers?

• The limiting resolution of a broad class of spectrometers is given (in wavenumbers) by the inverse of a characteristic dimension *L* of the spectrometer

Fourier-transform spectrometer



Grating spectrometer



 $\Delta \nu (\mathrm{res}) \approx 1/L$

- We use slow-light methods to design spectrometers with resolution that exceeds this conventional limit by a factor as large as the group index.
- This ability allows us to miniaturize spectrometers with no loss of resolution, for "lab-on-a-chip" applications.

Our Goal

Replace this:



with this:



Our Approach: Chip-Scale Slow-Light Spectrometer

- The spectral sensitivity of an interferometer is increased by a factor as large as the group index of a material placed within the interferometer.
- We want to exploit this effect to build chip-scale spectrometers with the same resoluation as large laboratory spectrometers



• We use line-defect waveguides in photonic crystals as our slow light mechanism

Slow-down factors of greater than 100 have been observed in such structures.

Shi, Boyd, Gauthier, and Dudley, Opt. Lett. 32, 915 (2007) Shi, Boyd, Camacho, Vudyasetu, and Howell, PRL. 99, 240801 (2007) Shi and Boyd, J. Opt. Soc. Am. B 25, C136 (2008).



Laboratory Characterization of the Slow-Light Mach-Zehnder Interferometer



Interference fringes



- Resolution (quarter wave) is
 17 pm or 2.1 GHz or 0.071 cm⁻¹
- (Slow-light waveguide is only 1 mm long!)

Magaña-Loaiza, Gao, Schulz, Awan, Upham, Dolgaleva, and Boyd, Opt. Lett. 41, 1431 (2016).

Slow Light and Fast Light: A Tutorial Review

- 1. Introduction to slow light
- 2. Mechanisms of slow light
- 3. Backwards light and causality
- 4. Slow light spectrometer
- 5. Optical forces
- 6. Fresnel drag
- 7. Other applications

Why Care about Optical Forces?

- Optical levitation
- Optical tweezers
- Optomechanical systems
- But can we *control* optical forces.

- Yes! Photon momentum and optical forces depend on both refractive index and group index of optical materials.

- The "slow light" community knows how to manipulate the group velocity of light.

Kinematic Properties of Slow and Fast Light

Poynting's Theorem when derived for a dispersive medium leads to the conclusion that

$$S = \frac{1}{2} n \epsilon_0 c E^2 \quad \text{(intensity)}$$
$$u = \frac{1}{2} n n_g \epsilon_0 E^2 \quad \text{(energy density)}$$

where

$$v_g = c/n_g$$
 (group velocity).

It thus follows that

$$S = u v_g.$$

Note:

Large enhancement of stored energy But no enhancement of E!

See, e.g., Haus, Landau and Lifshitz, Milonni, or Harris and Hau

In vacuum: $p = (\hbar \omega/c)$

Abraham form (for light in matter) $P = E \times H/c^2$ (EM momentum density) $p = (\hbar \omega/c)(1/n_g)$ (photon momentum)

Minkowski form (for light in matter) $P = D \times B$ (EM momentum density) $p = (\hbar\omega/c)(n^2/n_g)$ or $p = (\hbar\omega/c) n$ photon momentum

One way or other, photon momentum very small in slow-light medium

See, e.g., Barnett, PRL (2010), Milonni and Boyd, AOP (2010).

Einstein-Balazs Argument Supports the Abraham Form



- Argue that center of mass-energy must move with a constant velocity
- When photon wavepack enters block, it slows down. Block thus receives a kick into the forward direction.
- When photon leaves block, block receives backward kick and returns to rest.
- Block undergoes longitudinal displacement of

$$\Delta z = (n_g - 1)L\,\hbar\omega/(Mc^2)$$

• Simple kinematic argument shows that momentum of photon in block is

$$p = \hbar \omega / (n_g c)$$
 Abraham form!

Fermi's Argument Supports the Minkowski Form



photon in medium of refractive index $n(\omega)$

atom with mass m and resonance frequency ω_0

- Fermi describes Doppler effect in terms of atomic recoil (RMP, 1932)
- Atom can absorb only if $\omega pprox \omega_0 (1-nv/c)$
- Conservation of energy and momentum Initial energy $= \hbar \omega + \frac{1}{2}mv^2$ Final energy $= \hbar \omega_0 + \frac{1}{2}mv'^2$ Initial momentum = p + mv Final momentum = mv'
- Solve: find photon momentum *p* in medium given by

$$p = n \, \hbar \omega / c$$
 Minkowski form!

Which is Correct, Abraham or Minkowski?

PRL 104, 070401 (2010)

PHYSICAL REVIEW LETTERS

week ending 19 FEBRUARY 2010

S

Resolution of the Abraham-Minkowski Dilemma

Stephen M. Barnett

Department of Physics, SUPA, University of Strathclyde, Glasgow G4 0NG, United Kingdom (Received 7 October 2009; published 17 February 2010)

The dilemma of identifying the correct form for the momentum of light in a medium has run for a century and has been informed by many distinguished contributions, both theoretical and experimental. We show that *both* the Abraham and Minkowski forms of the momentum density are correct, with the former being the kinetic momentum and the latter the canonical momentum. This identification allows us to explain why the experiments supporting each of the rival momenta gave the results that they did. The inclusion of dispersion and absorption provides an interesting subtlety, but does not change our conclusion.

DOI: 10.1103/PhysRevLett.104.070401

PACS numbers: 03.50.De, 42.50.Nn, 42.50.Wk

 $\mathbf{p}_{kin}^{med} + \mathbf{p}_{Abr} = \mathbf{p}_{can}^{med} + \mathbf{p}_{Min}$

Total momentum (field plus material) the same in either treatment!

In vacuum: $p = (\hbar \omega / c)$

Abraham form (for matter)

 $P = E \times H/c^2$ (EM momentum density)

 $p = (\hbar \omega / c)(1/n_g)$ (photon momentum)

It is the kinetic (as in mv) momentum

It is the momentum of the field (alone)

It is what comes out of Balazs's moving block analysis

Minkowski form (for matter)

 $P = D \times B$ (EM momentum density)

 $p = (\hbar \omega/c)(n^2/n_g)$ or $p = (\hbar \omega/c) n$ photon momentum

It is the canonical momentum (as in $h/\lambda_{deBroglie}$)

It is the momentum of field and (at least part of that of the) matter It is what comes out of a Doppler shift analysis

One way or other, photon momentum very small in slow-light medium See, e.g., Barnett, PRL (2010), Milonni and Boyd, AOP (2010).

Slow Light and Fast Light: A Tutorial Review

- 1. Introduction to slow light
- 2. Mechanisms of slow light
- 3. Backwards light and causality
- 4. Slow light spectrometer
- 5. Optical forces
- 6. Fresnel drag
- 7. Other applications
Photon Drag Effects with Slow Light

We would like to use the dependence of the photon momentum on the group index as a means to control optical forces.

As a first step down this pathway, we are studying how to control photon drag effects using slow light.

The Velocity of Light in Moving Matter: Fresnel Drag (or Ether Drag) Effects

• Fizeau (1859): Longitudinal photon drag:

Velocity of light in flowing water.

V = 700 cm/sec; L = 150 cm; displacement of 0.5 fringe.



• Modern theory: relativistic addition of velocities

$$v = \frac{c/n + V}{1 + (V/c)(1/n)} \approx \frac{c}{n} + V\left(1 - \frac{1}{n^2}\right) - Fresnel "drag" coefficient$$

• But what about slow-light media?

Fresnel Drag in a Highly Dispersive Medium

Light Drag in a Slow Light Medium (Lorentz)

$$u \simeq \frac{c}{n} \pm v \left(1 - \frac{1}{n^2} + \frac{n_g - n}{n^2} \right)$$

We Use Rubidium as Our Slow Light Medium

• Transmission spectrum of Rb around D₂ transition:



• Group index of Rb around D_2 line at T=130



Safari, De Leon, Mirhosseini, Magana-Loaiza, and Boyd Phys. Rev. Lett. 116, 013601 (2016)



• Change in phase velocity is much larger than velocity of rubidium cell. Implications for new velocimeters?

Conclusions

• A maximum drag speed of 205 m/s was measured in a highly dispersive medium (hot Rb vapor).

• This effect is at least two orders of magnitude larger than that observed to date.

• Much larger dispersion can be achieved in Rb atoms using electromagnetically induced transparency (n_g as high as 10⁷).

Transverse Photon Drag



$$\Delta x = (vL/c)(n_g - 1/n_\phi)$$

For L = 25 mm, v = 2000 cm/s, displacement = 6 nm.

Measured by R.V. Jones, 1972.

Rotary Photon Drag: An image viewed through a spinning window



Theory says that transmitted image is rotated! (rotary photon drag)

(Polarization rotation measured earlier by Jones.)

Image rotation never previously observed (although implied by work of Leach et al., PRL 2008.)

Effect scales as group index!

Franke-Arnold, Gibson, Boyd and Padgett, Science, 2011

Observation of Rotary Photon Drag

The world as seen through a spinning window.

(Laser-excited ruby has a group index of 10^6 .)



Effect clearly visible by eye!



Franke-Arnold, Gibson, Boyd and Padget, Science, 2011

Slow Light and Fast Light: A Tutorial Review

- 1. Introduction to slow light
- 2. Mechanisms of slow light
- 3. Backwards light and causality
- 4. Slow light spectrometer
- 5. Optical forces
- 6. Fresnel drag
- 7. Other applications

Applications of Slow and Fast Light

- Buffers and regenerators for telecom
- Slow/fast light for interferometery
- Phased- and synchronized-array laser radar
- Construction of quantum memories

9. Spontaneous and Stimulated Light Scattering





complete destructive interference!

<u>Spontaneous light scattering</u> Fluctuations (e.g., in E) occur due to thermal (or quantum mechanical zero-point) excitation.

<u>Stimulated light scattering</u> Fluctuations are induced by the incident laser field.







Relation Between Spontaneous and Stimulated Light Scattering

$$\begin{array}{c} & \longrightarrow \overline{z} \\ & & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & & \\ & & & \\ & & & \\$$

 $\frac{dm_s}{dz} = Dm_L(m_s+1)$ # of laser photons per mode # of Stokes photons per mode

For
$$m_s \ll 1$$

 $m_s(z) = m_s(o) + Dm_L Z + M_s(z)$
linear increase with length of scattering medium
 \Rightarrow spontaneous scattering

For \underline{m}_{s} >>) $m_{s}(z) = m_{s}(o) \in Gz$ exponential growth \Rightarrow stimulated scattering

 $G = D m_{L}$

We can relate G to the spontaneous scattering cross section:

$$G = \frac{4\pi^{3}Nc^{2}}{\pi\omega_{L}\omega_{s}^{2}n_{s}^{2}} \left(\frac{\partial^{2}\sigma}{\partial\omega_{s}\partial\Omega}\right) I_{L}$$

(R.W. Hellwarth, Phys. Rev <u>130</u> 1850, 1963)





Positive feed back: laser + Stokes --- > sound laser + sound Stokes Frequencies: $W_2 = W_1 - \Omega_B \quad \text{where} \quad \Omega_{\overline{g}} = |\overline{g}| \mathcal{V} \quad \text{sound}$ $\overline{g} = \overline{k}_1 - \overline{k}_2 \simeq 2\overline{k}_1 \quad (\text{since } \mathcal{V} < c, \Omega < c W_1)$ $\Rightarrow \Omega_{B} = 2W \frac{V}{C/n} \qquad \left(\Omega_{B}/2\pi \approx 5 CH_{Z}\right)$

Coupling among waves : electrostriction
Material tends to be drawn
into regions of high field.

$$P_{st} = -Ve = \frac{E^2}{8\pi}$$

 $Ve = P \frac{2E}{2P} \approx 1$ for condense
matter.
 $\Delta P = -C P_{st}$ where $C = \frac{1}{P} \left(\frac{2P}{2P}\right) = \text{compress: bility}$
 $\Delta E = \left(P \frac{2E}{2P}\right) \frac{1}{P} \Delta P$

210

$$\frac{\text{Theory of SBS}}{\tilde{E}_{1}(z,t) = A_{1}(z,t) e^{i(R_{1}z-\omega_{1}t)} + tcc} = E_{1} \xrightarrow{z} E_{1}(z,t) = A_{1}(z,t) e^{i(R_{1}z-\omega_{1}t)} + tcc}$$

$$\tilde{E}_{2}(z,t) = A_{2}(z,t) e^{i(R_{1}z-\omega_{1}t)} + tcc} = \begin{cases} \overline{b}^{z}zR_{1} \\ \Omega \equiv \omega_{1}-\omega_{2} \end{cases}$$
density $\tilde{\rho}$ obeys a constic wave equation
$$\frac{\partial^{2}\tilde{\rho}}{\partial t^{2}} - \Gamma'\nabla^{2}\frac{\partial\tilde{\rho}}{\partial t} - \upsilon^{2}\nabla^{2}\tilde{\rho} = +\nabla\cdot\tilde{f}$$

$$\Gamma' = \frac{1}{\rho}\left[\frac{u}{3}\eta_{s} + \eta_{b} + \frac{\kappa}{c_{p}}(\gamma-1)\right]$$

$$\tilde{f} = \nabla\beta_{s} \qquad \beta_{s} + -\gamma_{e}\left[\frac{\langle \tilde{E}^{z}\rangle}{g\pi}\right]$$
Thus
$$\nabla\cdot\tilde{f} = \frac{\sqrt{e}}{4\pi}\left[A_{1}A_{2}^{*}e^{i(R_{2}-\Omega t)} + tcc\right]$$
In the SVAA, acoustic wave equation becomes
$$-2i\Omega\frac{\partial\rho}{\partial t} + \left(\Omega_{R}^{2}-\Omega^{2}-i\Omega\Gamma_{R}\right)\rho - 2ig\upsilon^{2}\frac{\partial\rho}{\partial z}$$

$$= \frac{\sqrt{e}}{4\pi}\frac{g^{2}}{A_{1}}A_{2}^{*}$$
where
$$\Gamma_{R} = g^{2}\Gamma' \qquad \Omega_{R} = 2\omega\frac{\tau}{c/n}$$

. .

250

11

$$\frac{\text{Theory of SBS}}{\text{Simplifications}: (1) \ \partial P/\partial t = 0} \quad \text{in steady state}$$

$$(2) \ \partial P/\partial z = 0 \quad (\text{phonons strongly damped})$$

$$(ds \leq 10 \text{ Jum})$$

$$P(z, t) = \frac{\text{Yeg}^2}{4\pi} \quad \frac{A_1 A_2^*}{\Omega_B^2 - \Omega^2 - i \Omega \Gamma_B}$$
Optical field obeys

$$\frac{\partial^2 \widetilde{E}_i}{\partial z^2} = \frac{1}{(c/n)^2} \frac{\partial^2 \widetilde{E}_i}{\partial t^2} = \frac{4\pi}{c^2} \frac{\partial^2 \widetilde{P}_i}{\partial t^2}$$

where

$$P^{NL} = \Delta X \tilde{E} = \frac{\Delta E}{4\pi} \tilde{E} = \frac{1}{4\pi \rho_{o}} \left(\rho \frac{\partial E}{\partial \rho} \right) \left(\tilde{\rho} - \rho_{o} \right) \tilde{E}$$
$$= \frac{1}{4\pi \rho_{o}} Y_{e} \left(\tilde{\rho} - \rho_{o} \right) \tilde{E}$$

Thus

$$\frac{\partial A_{1}}{\partial z} + \frac{1}{c/n} \frac{\partial A_{1}}{\partial t} = \frac{i\omega Y_{e}}{2nc\rho_{o}} \rho A_{2}$$

$$-\frac{\partial A_{2}}{\partial z} + \frac{1}{c/n} \frac{\partial A_{2}}{\partial t} = \frac{i\omega Y_{e}}{2nc\rho_{o}} \rho^{*} A_{1}$$
At the Brillowin resonance $(\Omega = \Omega_{B})$ and in steady state
$$\frac{dA_{1}}{dz} = -\frac{\omega}{8\pi} \frac{Y_{e}^{2} g^{2}}{8\pi} (A_{2} A_{2}^{*}) A_{1}$$

$$\frac{dA_{2}}{dz} = -\frac{\omega}{8\pi} \frac{Y_{e}^{2} g^{2}}{8\pi} (A_{1} A_{1}^{*}) A_{2}$$

-

-

.*****____.





We can predict the Stokes output intensity :



$$T_{s}(L) = T_{s}(o) e^{gT_{L}}L$$



i ypicai Kaman meura		
material	Δν (cm ⁻¹)	gain g x 10^3 (cm/MW)
LIQUIDS		
benzene	992	3
water	3290	0.14
N2	2326	17
O ₂	1555	16
GASES		
methane	2916	0.66 (10 atm, 500 nm)
hydrogen	4155 (vibrational)	1.5 (above 10 atm)
	450 (rotational)	0.5 (above 0.5 atm)
deuterium	2991 (vibrational)	1.1 (above 10 atm)
N ₂	2326	0.071 (10 atm, 500 nm)
O ₂	1555	0.016 (10 atm, 500 nm)

Typical Raman media

U. Simon and F. K. Tittel, Nonlinear Optical Frequency Conversion Techniques, in *Methods of Experimental Physics, Vol. III (Lasers and Optical Devices)*, R. G. Hulet, and F. B. Dunning, eds., Academic Press, 1994.







Stimulated Scattering Processes

Stimulated Royleigh Scattering interaction with nonpropagating density (temperature) waves (coupling can be electrostrictive of thermal) $W_{L} \longrightarrow W_{S} = W_{L} - \Omega \longrightarrow W_{S}$ $W_{S} = W_{L} - \Omega \longrightarrow W_{S}$ $W_{S} = W_{L} - \Omega \longrightarrow W_{S}$

<u>Stimulated Raman Scattering</u> interaction with vibrational degree of freedra

Stimulated Rayleigh-Wing Scattering interaction with orientational degree of freedom $\sim 2ps_{7}$ $W_{L} \longrightarrow I \times I \times I \times I$ $W_{S} = W_{L} - \Omega \longrightarrow I / Torientation$ aligned molecules randomly oriented molecules



Experiment in Self Assembly



Joe Davis, MIT

Thank you for your attention!

